

First Direct Seeding at 38 nm

Christoph Lechner

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on behalf of the sFLASH team



Universität Hamburg
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Bundesministerium
für Bildung
und Forschung

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Outline

- Introduction / Motivation
- The experiment
 - Layout
 - Procedures
- Seeding results at $\lambda = 38 \text{ nm}$

Motivation for Seeding

SASE Free-electron lasers

- pulsed radiation with wavelengths down to sub-Å
- transverse coherence
- typically limited longitudinal coherence: multiple uncorrelated longitudinal modes present

with seeding:

- amplify external coherent optical field
- longitudinal coherence of the FEL pulse determined by the external field

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Goals:

- high peak power (GW level)
- stable pulse spectrum and energy
- high longitudinal coherence

FLASH at DESY (Hamburg)

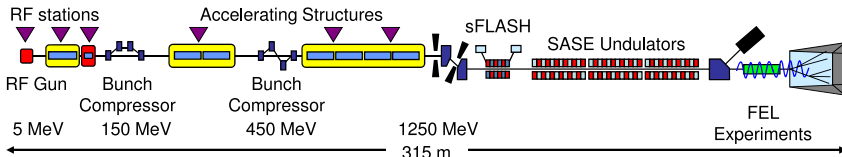
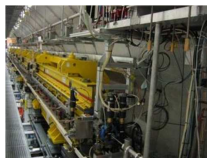
Normal conducting 1.3 GHz RF gun



TESLA type superconducting accelerating modules (1.3 GHz)



Fixed gap SASE undulators



3rd harmonic module (3.9 GHz)



Diagnostics and matching

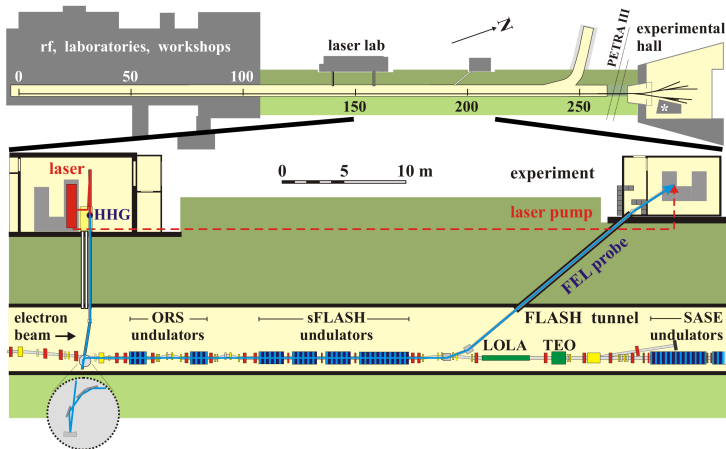


Variable gap sFLASH undulators



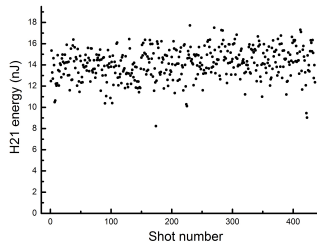
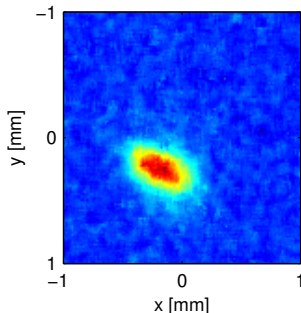
FEL Experimental Hall

Layout of the sFLASH Experiment



HHG Source Performance

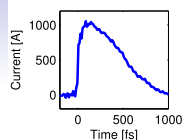
Typical HHG seed beam profile at the entrance of the first undulator:



- ✓ At the source: energy in 21st harmonic 14 nJ (10% rms stability)

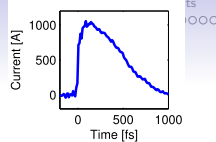
Experimental Procedure

- setup accelerator for 700 MeV
 - bunch charge 0.5 nC
 - feedback systems for compression and energy
- establish high FEL gain at correct wavelength
 - tuning sFLASH to SASE
 - spectral overlap of 21st harmonic ($\lambda = 38.1$ nm) and sFLASH SASE



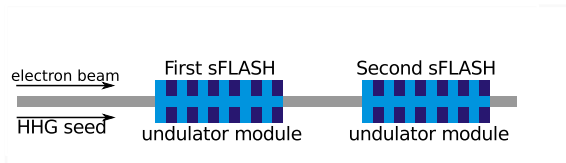
Current profile reconstructed
from a THz single-shot spectrum

Experimental Procedure

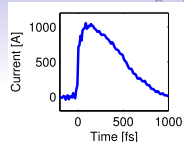


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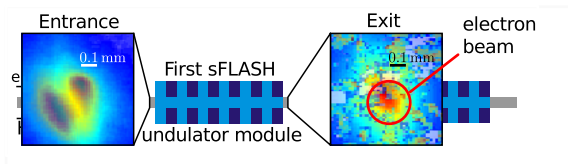


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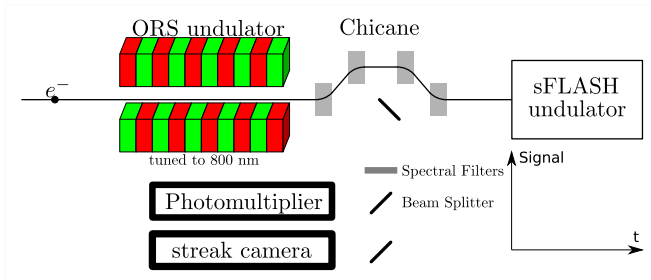
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typical transverse profiles (electrons and HHG) at the first sFLASH undulator

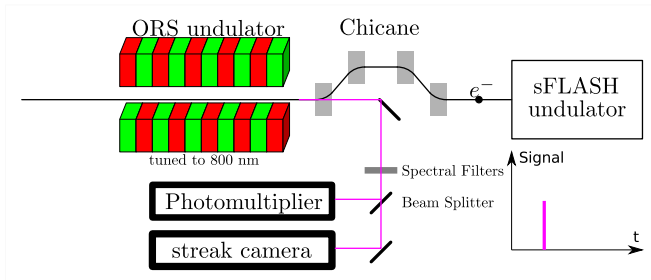
Experimental Procedure

- temporal overlap
 - down to 1 ns: photomultiplier + oscilloscope
 - down to 10 ps: streak camera
 - finally: time scan (100 fs steps)



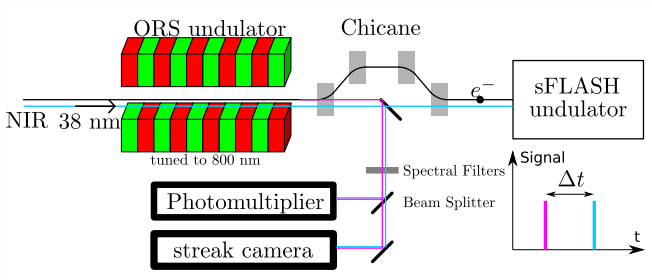
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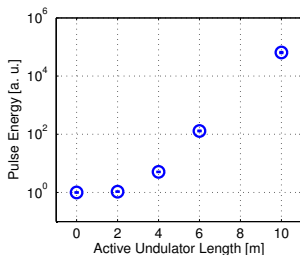
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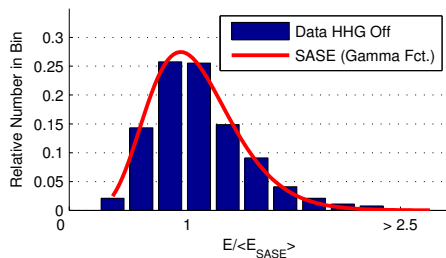


FEL Amplifier Characterization

SASE mode



Power gain length about
0.65 m

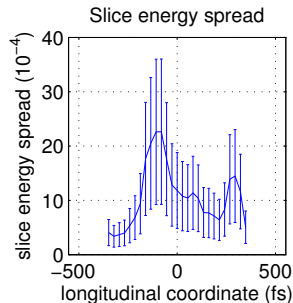
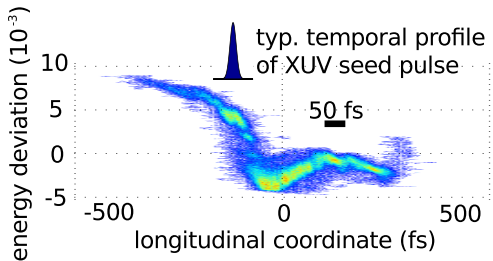


Number of longitudinal modes:
 $M = 8.3 \pm 1.2$

⇒ Coherence time 6 fs, radiation pulse length ~ 50 fs

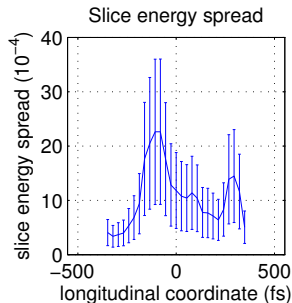
LOLA TDS

With the LOLA transverse deflecting structure (TDS) one can measure the longitudinal phase space after sFLASH undulators ...



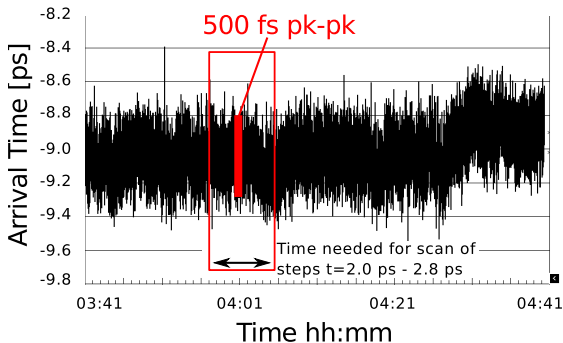
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Electron Bunch Arrival Time During Measurements

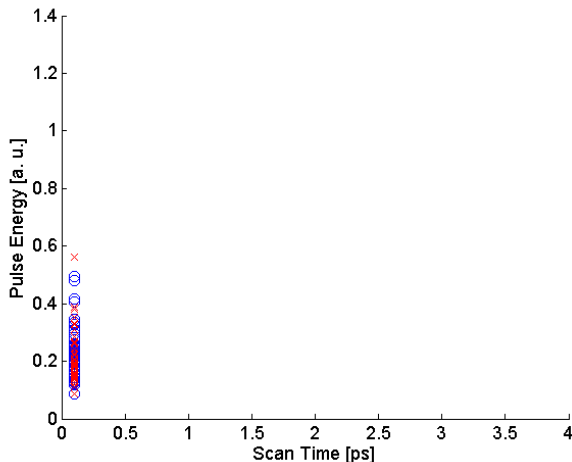
- arrival time feedback off due to single bunch operation
- future sFLASH scenario: make use of existing arrival time feedbacks in multibunch operation mode



SASE pulse duration estimated to 50 fs
500 fs pk-pk electron arrival jitter
⇒ limits probability for temporal overlap

Temporal Fine Scan

Transverse, wavelength overlap and coarse temporal overlap established.
Scan time to establish overlap.



○
HHG source
on

×
HHG source
off

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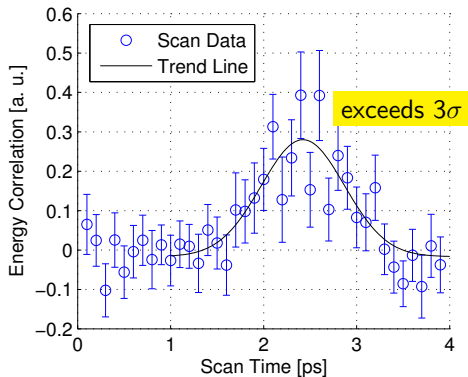


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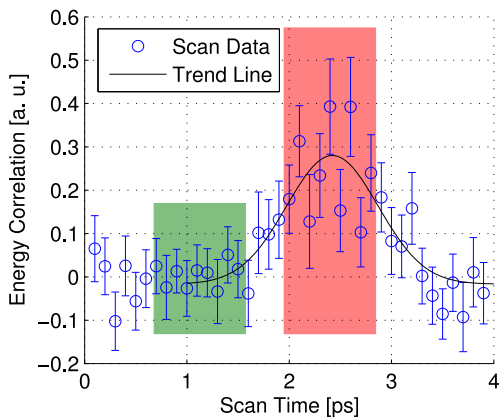
HHG source
off

FEL Pulse Energy vs. Time Offset



For each scan step: Correlation of XUV seed pulse energy at the source and pulse energy in the FEL pulse

Alternative Data Analysis

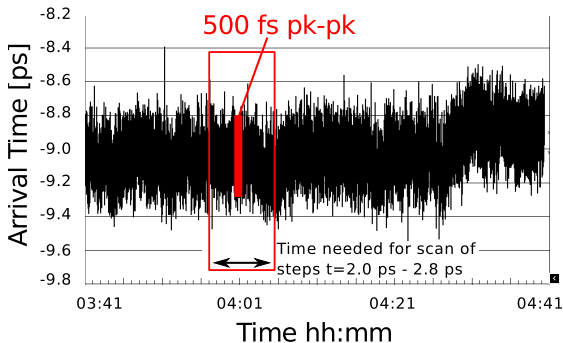


Alternative Data Analysis

Histograms removed for copyright reasons (paper submitted).

Electron Bunch Arrival Time During Measurements

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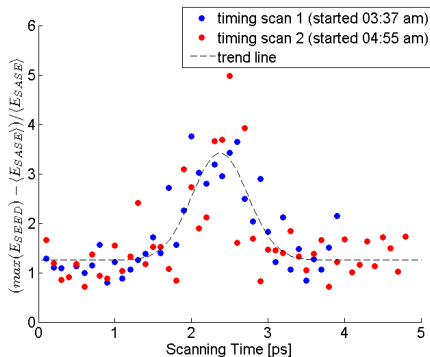


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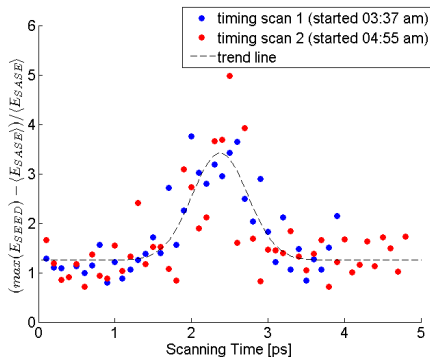
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Energy Contrast Measurement



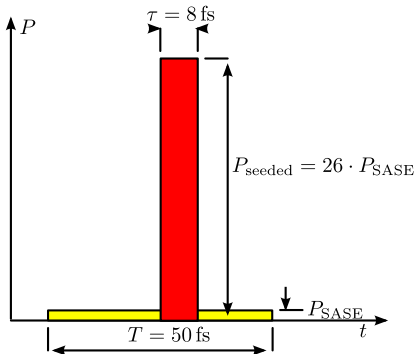
Energy Contrast Measurement



What can we conclude from $\frac{\max(E_{\text{seed}}) - \langle E_{\text{SASE}} \rangle}{\langle E_{\text{SASE}} \rangle} = 4$ for excess of instantaneous power in the seeded part of the pulse?

From Energy Contrast to Power Increase

Model for expected time profile of the photon pulse



Assumptions:

- energy contrast of 4
- SASE pulse length $T = 50 \text{ fs}$
- HHG pulse length $\tau = 8 \text{ fs}$

Linear FEL amplifier

$$\frac{E_{\text{tot}}}{E_{\text{SASE}}} = \frac{T \cdot P_{\text{SASE}} + \tau(P_{\text{seeded}} - P_{\text{SASE}})}{T \cdot P_{\text{SASE}}}$$

$$\Rightarrow P_{\text{seeded}} = 26 \cdot P_{\text{SASE}}$$

Summary and Outlook

- first seeding at wavelength $\lambda = 38$ nm demonstrated
 - power contrast approx. 30
- lessons learned:
 - need for HHG pulse characterization at the entrance of the undulator
 - stability of temporal overlap limits fraction of seeded bunches
 - feedbacks have to be applied to electron beam and photon beam parameters

Summary and Outlook

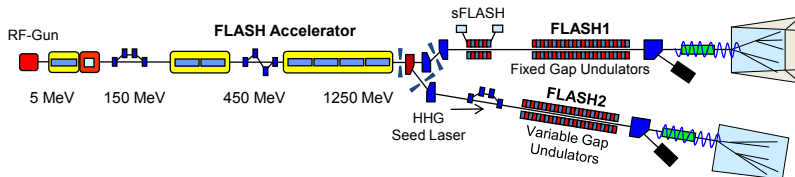
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Outlook:

- spectral phase control
- temporal characterization of seeded FEL radiation
- towards shorter wavelengths
- run sFLASH parallel to FLASH SASE

FLASH2 Beamline

Seeding will be a key asset of FLASH2.



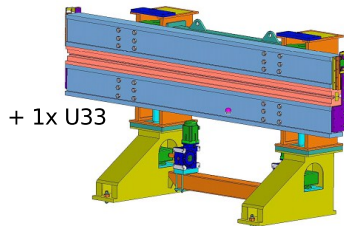
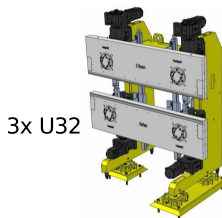
Posters:

- Sven Ackermann et al: “Optimization of HHG Seeding between 10 nm to 40 nm”, TUPD11
- Katja Honkavaara: “Status of the FLASH II Project”, WEPD07

Thank you for your attention

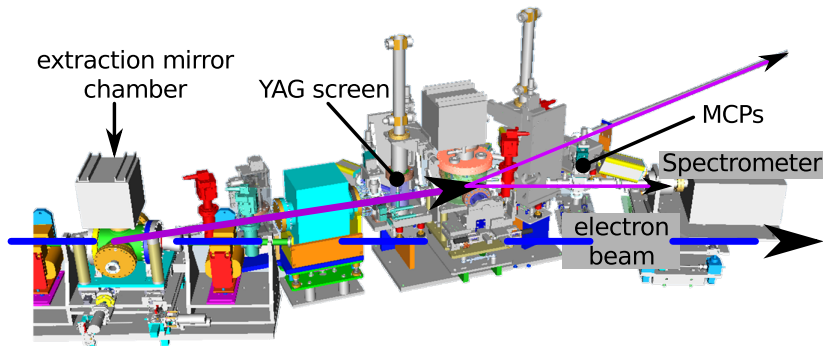
S. Ackermann, A. Azima, S. Bajt, J. Bödewadt, F. Curbis, H. Dachraoui, H. Delsim-Hashemi, M. Drescher, S. Düsterer, B. Faatz, E. Hass, U. Hipp, K. Honkavaara, R. Ischebeck, S. Khan, T. Laarmann, C. Lechner, T. Maltezopoulos, V. Miltchev, M. Mittenzwey, M. Rehders, J. Rönsch-Schulenburg, J. Roßbach, H. Schlarb, S. Schreiber, L. Schroedter, R. Tarkeshian, M. Tischer, V. Wacker, M. Wieland

sFLASH Undulator Modules



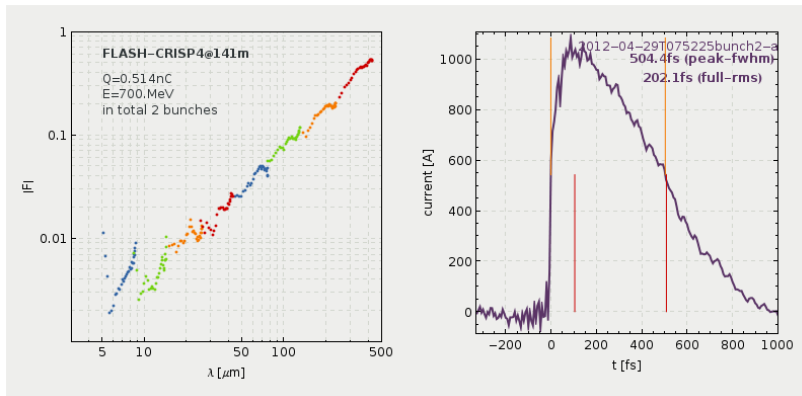
	U32	U33
Minimum gap [mm]	9.0	9.8
Period length [mm]	31.4	33
No. of poles	120	240
Length [m]	2	4
maximum K value	2.72	3.03

Photon Extraction and Diagnostics



- Located after the sFLASH undulators
- Micro-channel plates (MCPs) used to detect FEL radiation pulse energy

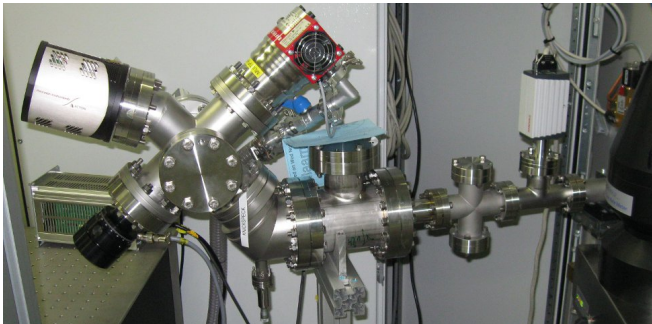
CRISP4



Left: single-shot THz spectrum

Right: Reconstructed current profile

Spectral Overlap



Spectral Overlap

